# Systematic calculation of photoresponse with the Skyrme functional 

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## Time-dependent DFT

Time-dependent Kohn-Sham (TDKS) equation
Time-dependent Hartree-Fock (TDHF) equation

$$
\begin{gathered}
i \frac{\partial}{\partial t} \phi_{i}(t)=\left\{h(t)+V_{\mathrm{ext}}(t)\right\} \phi_{i}(t) \\
i \frac{\partial}{\partial t} \rho(t)=\left[h(t)+V_{\mathrm{ext}}(t), \rho(t)\right] \\
\rho(\vec{r}, t)=\sum_{i=1}^{N}\left|\phi_{i}(\vec{r}, t)\right|^{2} \\
h(t)=h[\rho(t)]
\end{gathered}
$$

## TDDFT for superfluid systems

Time-dependent Kohn-Sham-Bogoliubov (TDKSB) equation Time-dependent Hartree-Fock-Bogoliubov (TDHFB) equation

$$
\begin{aligned}
& i \frac{\partial}{\partial t} \Psi_{i}(t)=\left\{H(t)+V_{\mathrm{ext}}(t)\right\} \Psi_{i}(t) \\
& i \frac{\partial}{\partial t} R(t)=\left[H(t)+V_{\mathrm{ext}}(t), R(t)\right]
\end{aligned}
$$

$$
\begin{aligned}
& \Psi_{i}=\binom{U_{i}}{V_{i}} \quad H(t)=H[R(t)]=\left(\begin{array}{cc}
h & \Delta \\
-\Delta^{*} & -h^{*}
\end{array}\right) \\
& R(t)=\sum_{i} \Psi_{i} \Psi_{i}^{+}=\left(\begin{array}{cc}
\rho(t) & \kappa(t) \\
-\kappa^{*}(t) & 1-\rho^{*}(t)
\end{array}\right)
\end{aligned}
$$

## Small-amplitude approximation --- Linear response (RPA) equation ---

$$
\begin{aligned}
& \left\{\left(\begin{array}{cc}
A & B \\
B^{*} & A^{*}
\end{array}\right)-\omega\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\right\}\binom{X_{m i}(\omega)}{Y_{m i}(\omega)}=-\binom{\left(V_{\text {ext }}\right)_{m i}}{\left(V_{\text {ext }}\right)_{i m}} \\
& A_{m i, n j}=\left(\varepsilon_{m}-\varepsilon\right) \delta_{m n} \delta_{i j}+\left\langle\phi_{m}\right| \frac{\partial h}{\partial \rho_{n j}}\left|\phi_{\rho_{0}}\right\rangle \\
& B_{m i, n j}=\left\langle\phi_{m} \frac{\partial h}{\partial \rho_{j n}}\right| \rho_{\rho_{0}}\left|\phi_{i}\right\rangle \\
& \text { - Tedious calculation of residual interactions } \\
& \text { - Computationally very demanding, } \\
& \text { especially for deformed systems. }
\end{aligned}
$$

However, in principle, the self-consistent single-particle Hamiltonian should contain everything. We can avoid explicit calculation of residual interactions.

## Finite Amplitude Method

T.N., Inakura, Yabana, PRC76 (2007) 024318.

Residual fields can be estimated by the finite difference method:

$$
\begin{aligned}
& \delta h(\omega)=\frac{1}{\eta}\left(h\left[\left\langle\psi^{\prime}\right|,|\psi\rangle\right]-h_{0}\right) \\
& \left|\psi_{i}\right\rangle=\left|\phi_{i}\right\rangle+\eta\left|X_{i}(\omega)\right\rangle, \quad\left\langle\psi_{i}^{\prime}\right|=\left\langle\phi_{i}\right|+\eta\left\langle Y_{i}(\omega)\right|
\end{aligned}
$$

Starting from initial amplitudes $X^{(0)}$ and $Y^{(0)}$, one can use an iterative method to solve the following linear-response equations.

$$
\begin{aligned}
& \omega\left|X_{i}(\omega)\right\rangle=\left(h_{0}-\varepsilon_{i}\right)\left|X_{i}(\omega)\right\rangle+\hat{Q}\left\{\delta h(\omega)+V_{\mathrm{ext}}(\omega)\right\}\left|\phi_{i}\right\rangle \\
& \omega\left\langle Y_{i}(\omega)\right|=-\left\langle Y_{i}(\omega)\right|\left(h_{0}-\varepsilon_{i}\right)-\left\langle\phi_{i}\right|\left\{\delta h(\omega)+V_{\mathrm{ext}}(\omega)\right\} \hat{Q}
\end{aligned}
$$

Programming of the RPA code becomes very much trivial, because we only need calculation of the single-particle potential, with different bras and kets.

## Finite amplitude method for superfluid systems

Avogadro and TN, PRC in press (arXiv:1104.3692)
Residual fields can be calculated by

$$
\begin{aligned}
& \delta h(\omega)=\frac{1}{\eta}\left\{h\left[\widetilde{v}^{*}, v\right]-h_{0}\right\} \\
& \left.\delta \Delta(\omega)=\frac{1}{\eta}\left\{\Delta \widetilde{v}^{*}, u\right]-\Delta_{0}\right\}
\end{aligned}
$$

$$
\begin{array}{ll}
v=V+\eta U X, & \widetilde{v}^{*}=V+\eta U^{*} Y \\
u=U+\eta V^{*} Y, & \widetilde{v}^{*}=U^{*}+\eta V^{*} X
\end{array}
$$

QRPA equations are

$$
\begin{array}{ll}
\left(E_{\mu}+E_{v}-\omega\right) X_{\mu \nu}+\delta H_{\mu \nu}=0 \\
\left(E_{\mu}+E_{\nu}+\omega\right) Y_{\mu \nu}+\delta \widetilde{H}_{v \mu}^{*}=0
\end{array} \quad\left(\begin{array}{cc}
\delta H_{\mu \nu} \\
\delta \widetilde{H}_{\mu \nu} &
\end{array}\right)=W^{+}\left(\begin{array}{cc}
\delta h & \delta \Delta \\
\delta \widetilde{\Delta}^{+} & -\delta h^{+}
\end{array}\right) W
$$

## Numerical Details

- SkM* interaction (no pairing)
-3D mesh in adaptive coordinate
- $\mathrm{R}_{\text {box }}=15 \mathrm{fm}$
- Complex energy with $\Gamma=1.0 \mathrm{MeV}$
- $\Delta \mathrm{E}=0.3 \mathrm{MeV}$ up to $\mathrm{E}=38.1 \mathrm{MeV}$ (128 points)
- Energy-paralleled calc. on PACS-CS

adaptive coordinate PRC71, 024301


PACS-CS @ Univ. of Tsukuba

## Electric Dipole strength distributions

- 3D mesh
- SkM*
- $\mathrm{R}_{\mathrm{box}}=15 \mathrm{fm}$
- No pairing

| $\frac{\operatorname{Sn}}{C d}$ |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |
| RuNu |  |  |  |  |  |  |  |  |
| MownN |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |

Ge 1


Fennn (NANMNNMNMNA Fe

Ti
40 spherical nuclei 176 prolate nuclei 59 oblate nuclei 72 triaxial nuclei

## Magic numbers for PDR emergence



## Next magic number: $\mathrm{N}=51$



## Magic numbers and low-/ orbits

- Magic numbers: $\mathrm{N}=15,29,51, \ldots$
- Importance of weakly bound orbits with $l=0,1$, and 2.



## PDR strength is correlated with any quantity?



Linear correlation was found for $\mathrm{R}_{\mathrm{n}}-\mathrm{R}_{\mathrm{p}}$ for neutron-deficient Sn (spherical) isotopes


## Universal correlation with skin thickness

- PDR fraction/ $\Delta R_{n p}$ shows a universal rate.
- The rate is about $0.2 / \mathrm{fm}$.



## Axially deformed superfluid nuclei

$$
\begin{aligned}
& \left(\begin{array}{cc}
h-\lambda & \Delta \\
-\Delta^{*} & -(h-\lambda)^{*}
\end{array}\right)\binom{U_{\mu}(\rho, z ; \sigma)}{V_{\mu}(\rho, z ; \sigma)}=E_{\mu}\binom{U_{\mu}(\rho, z ; \sigma)}{V_{\mu}(\rho, z ; \sigma)} z+\cdots \\
& \left\{\left(\begin{array}{cc}
A & B \\
B^{*} & A^{*}
\end{array}\right)-\omega\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\right\}\binom{X_{m i}(\omega)}{Y_{m i}(\omega)}=0 \underset{\rho}{\square}
\end{aligned}
$$

- HFB equations are solved in the 2D coordinate space, assuming the axial symmetry for the $\mathrm{SkM}^{*}$ functional with the cutoff of $E_{\mathrm{qp}}<60 \mathrm{MeV}$.
- The pairing energy functional is the one determined by a global fitting to deformed nuclei (Yamagami, Shimizu, TN, PRC 80, 064301 (2009))
- QRPA matrix is calculated in the quasiparticle basis ( $E_{2 q p}<60 \mathrm{MeV}$ ).
- Neglect the residual Coulomb interaction


## Shape phase transition in the EDF approach



## QRPA calculation of photoabsorption cross section

SkM* functional


## PDR in rare-earth nuclei



- Larger PDR strength for deformed nuclei
- Experimental data suggest a concentrated E1 strength in $\mathrm{E}=5.5-8 \mathrm{MeV}$.
- Calculation beyond QRPA is necessary.


## Summary

- Linearized (small-amplitude) TDDFT can be formulated in several different ways.
- Finite amplitude method (PRC76, 024318) provides an alternative approach to (Q)RPA.
- FAM does not require explicit calculations of residual interactions, thus, fully self-consistent calculations for deformed nuclei can be easily achieved.
- Systematic calculations of photoabsorption cross sections in light to heavy nuclei
- Reproduce the GDR peak and shape evolution
- Magic numbers for PDR ( $\mathrm{N}=15,29,51, \ldots$ )
- Universal correlation between the PDR fraction and the neutron skin thickness; $m_{1}(P D R) / m_{1} \approx(0.2 / f m) \Delta R_{n p}$.

