

### Nuclear Diagnostics at the National Ignition Facility

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(et. a lot of al.)

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# NIF is the culmination of a long line of glass laser systems developed at LLNL



### NIF concentrates all 192 laser beam energy in a football stadium-sized facility into a mm<sup>3</sup>

Matter Temperature >10<sup>8</sup> K Radiation Temperature >3.5 x 10<sup>6</sup> K Densities >10<sup>3</sup> g/cm<sup>3</sup> Pressures >10<sup>11</sup> atm

### Target Chamber June 1999

22EIM/bc + NIF-0609-16535a1



### ...In the target chamber



### **How NIF/ICF Works**



### "Layered-cryo" w/ hohlraum (indirect drive) vs. "Exploding pusher" (direct drive)

#### "Layered-Cryo"

- Laser energy produces ~300 eV x-rays in hohlraum "can," heating CH or Be capsule wall
- Cryogenic DT
   "layered" fuel shell
   with gas interior
- "Hot spot" ignites
   high ρR layer burn
- Yield up to 10<sup>19</sup> n





Cryogenic, x-ray driven, layered targets





# Lofty goal of the National Ignition Campaign (NIC): "



### Pitfalls to achieving ignition





### Neutron Spectra Diagnostics must span 9 orders of magnitude

NIF



All neutrons born within <100 ps.

#### Activation: Zr Neutron activation (NAD) measures yield for DT shots to absolute accuracy of $\pm$ 7%







NIF <u>MRS:</u> The MRS has been designed and implemented for simultaneous measurements of  $\rho R$ ,  $Y_{1n}$  and  $T_{ion}$ **MRS Spectrum THD-3 Med-Res** Low-Res 10<sup>4</sup> N11020  $t_{\rm f}$  = 125 µm  $t_f = 275 \ \mu m$ **Primaries Yield** 10  $A_a = 20 \text{ cm}^2$ Me< n DS-n d+ Counts / Magnet 10<sup>2</sup> **Deconvolved** TT-n width = T<sub>ion</sub>  $R_{f}$  = 26 cm 10 **Downscatter from 10-12MeV** neutrons *R<sub>a</sub>* = 570 cm 10<sup>0</sup> 10 15 0 5 Deuteron energy [MeV] **CR-39** Accuracy requirement for the MRS absolute yield measurement < 10% for  $Y_{1n}$  > 10<sup>14</sup>



#### <u>GRH:</u> Gamma Reaction History (GRH) measures Bang Time (w/in 30 ps) and Burn Width (w/in 15 ps) with Gas Cherenkov Detectors



- Bangtime agrees with Ntof\_BT data to within 100 ps
- Energy threshold of each cell set by gas pressure /composition
- 3-10MeV GRH will be fielded on DT implosions for yield and 4.4MeV carbon γ (ablator density measurement)



NIF-0211-20965.ppt



# **<u>nToF</u>**: lon temperatures, yields, and ρR from ~6 nToF detectors are calculated by an iterative process





## Neutron Imaging system has begun performance qualification process using exploding pusher shots



# First radiochemistry (RAGS) diagnostic utilizes $^{124}$ Xe(n, $\gamma$ ) and $^{124}$ Xe(n,2n) to measure average $\rho R$



NIF

Sandia National Laboratories



# Where we are now and where we're going

- Primary purpose of these first two diagnostics phases are to achieve ignition
- We are now (≈3 weeks ago) beginning to think about science-enabling diagnostics, including:
  - 1. Solid-debris collection (fast and slow)
  - 2. Energy resolving  $\gamma$ -ray detectors (bent crystal)
  - Fission-based low-energy neutron spectrometers (Supplements what Lee talked about)

...

# Coming up with an idea for a new diagnostic is a great way to get involved

### **Nuclear Physics AT NIF** (thanks Lee!) Nuclear Physics FOR NIF

- D-T fusion 16.7 MeV γ-ray branching ratio
- T-T neutron spectrum (<sup>6</sup>He breakup)
  - Sequential, di-neutron, or two-body?
- Nuclear-plasma interactions/rates/thermal population
  NEEC, NEET, etc.
- Reactions on highly-excited states
- Cross sections: (n,x) for radchem

# NIF is providing opportunities to explore new areas of nuclear physics





### **Nuclear physics needs**

#### T-T neutron spectrum at NIF-relevant energies (~10keV)





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### Nuclear Cross Sections for Charged Particles at Energies Relevant to Astrophysics are Difficult to Measure



By measuring reaction products at NIF the relevant cross sections are inferred

The achievement of ignition will provide unique research opportunities in astrophysics, stewardship physics, and inertial fusion energy studies



M. Junker et al., PRC 57, 2700 (1998)

## • Strongly screened reactions are relevant to stellar evolution



- First hit gives excited nuclear state
- Reactions from excited states, relevant to r-process nucleosynthesis of heavy elements
- Second hit reaction cross section uncertain
  - S. Libby, IFSA proceedings (2004)