

# Level Densities and $\gamma$ Strength Functions in $^{116,118,119}\text{Sn}$

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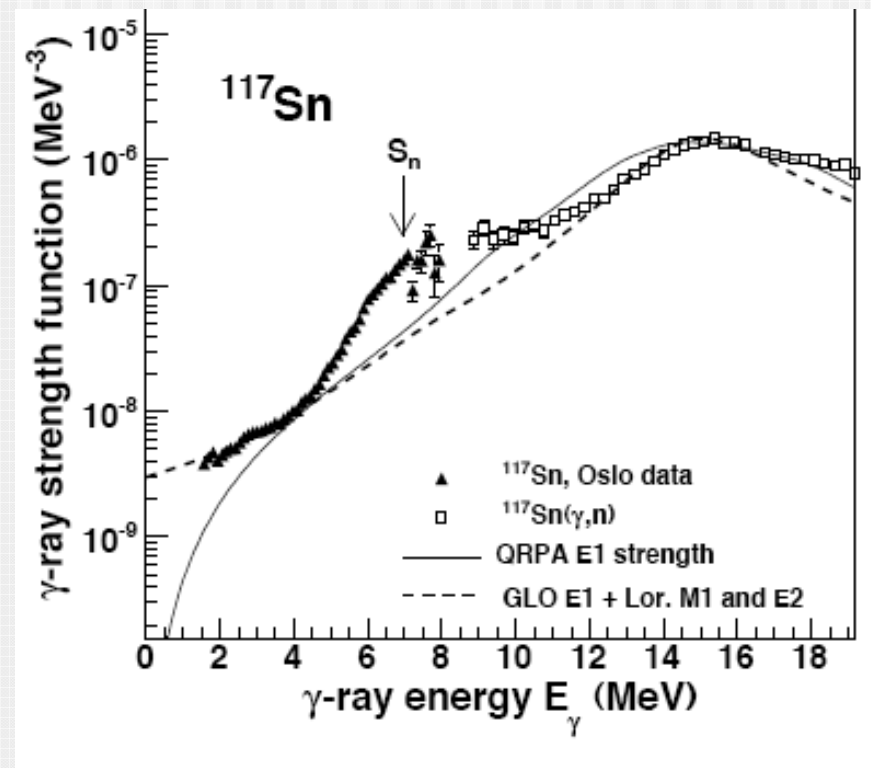
# Outline

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1. Experiment motivation
2. The Oslo method
3. Preliminary level densities and  $\gamma$  strength functions
  1.  $^{119}\text{Sn}$
  2.  $^{116,118}\text{Sn}$

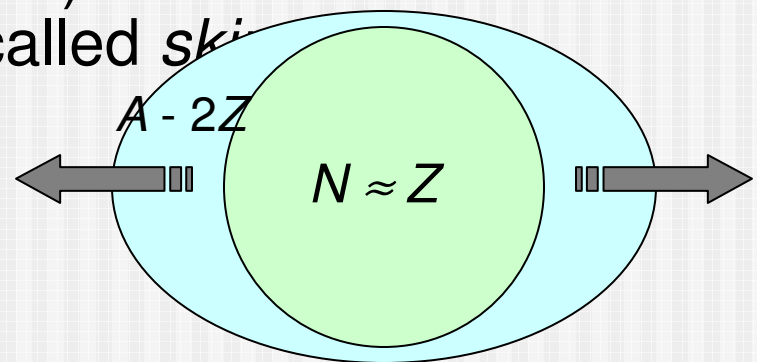
# Discovery of a new resonance in another Sn isotope

- $^{117}\text{Sn}$ : New small resonance in  $\gamma$  strength function.
- Recent PRL publication from Oslo.  
(*PRL* **102**, 162504 (2009).)
- “Pygmy” resonance.
- Located on the GEDR tail, for low  $E_\gamma$ , but higher than for the scissor mode.
- Peak  $E_\gamma \approx 9$  MeV.



# Possibilities for pygmy $\diamond$ 's origin

- Already theory predictions of small resonances in this area:
  - GMDR (M1 mode) - but strength too small;
  - Neutron-skin oscillation resonance (E1 mode).  
(*van Isacker et al., PRC 45, R13 (1992).*)
- Visualization of skin oscillation:
  - A non-moving core of  $Z$  protons and  $N \approx Z$  neutrons.
  - Excess neutrons ( $\approx A - 2Z$ ) oscillate in the nucleus  $\diamond$  outer part, so-called *skin*.

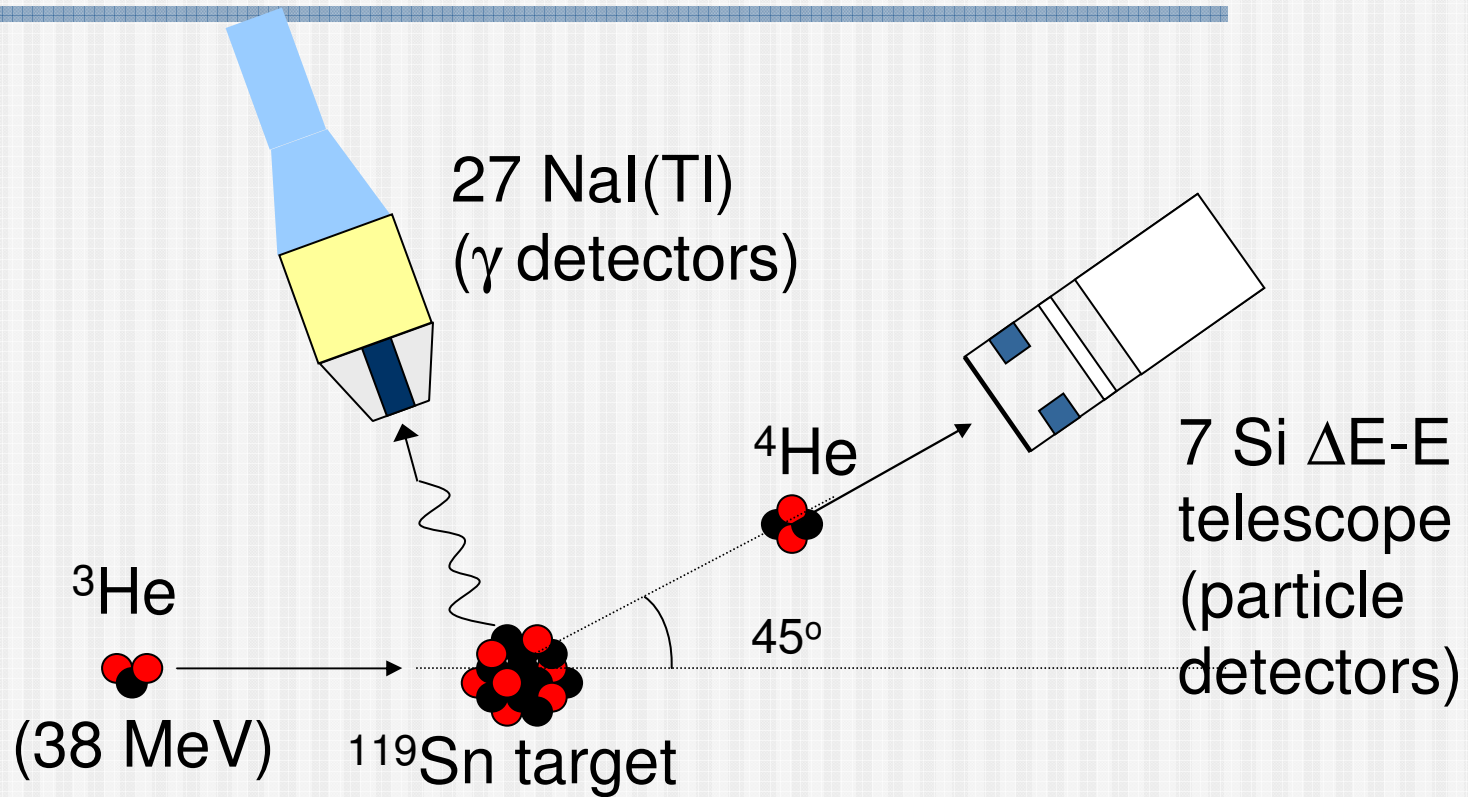


# Motivation for performing $^{118,119}\text{Sn}$ experiment

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- Confirm pygmy recently published in PRL.
- Study pygmies of different tin isotopes.

# Experimental setup



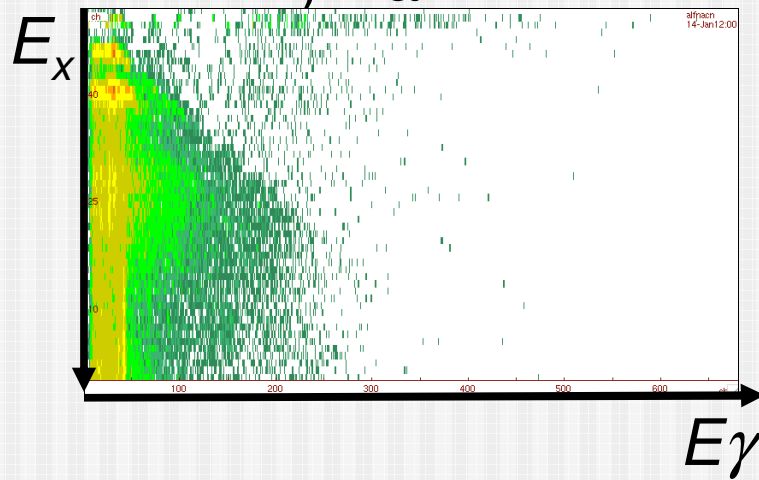


# The Oslo method

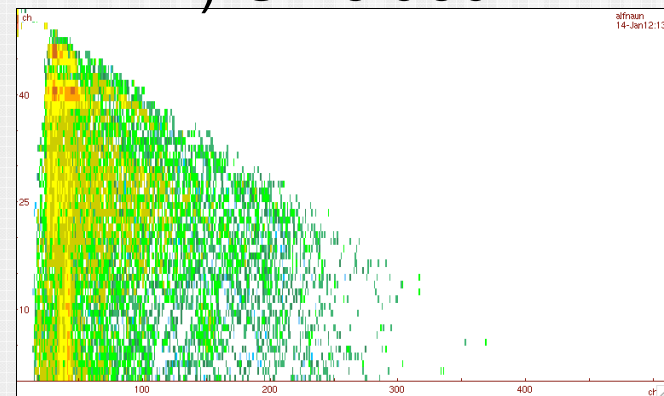
- Experimental detection of particle and  $\gamma$  coincidences,  $E_\gamma < B_n$ :
  - Pick-up reaction:  $^{119}\text{Sn} (^3\text{He}, ^4\text{He} \gamma) ^{118}\text{Sn}$ .
  - Inelastic scattering:  $^{119}\text{Sn} (^3\text{He}, ^3\text{He}' \gamma) ^{119}\text{Sn}$ .
- Unfolding of  $\gamma$  spectra using detector response function.
- Extraction of first generation  $\gamma$  s.
- First generation coincidence matrix  $(E_\gamma, E_x)$ .
- Simultaneously extract estimates for level density  $\rho$ , and for  $\gamma$  transmission coefficient  $T$ :
$$P(E_x, E_\gamma) \propto \rho(E_x - E_\gamma) \cdot T(E_\gamma).$$
- Determine  $\gamma$  strength function  $f(E_\gamma)$ :
$$f(E_\gamma) = T(E_\gamma) \cdot E_\gamma^{2L+1}.$$

# Matrix ( $E_\gamma$ , $E_x$ )

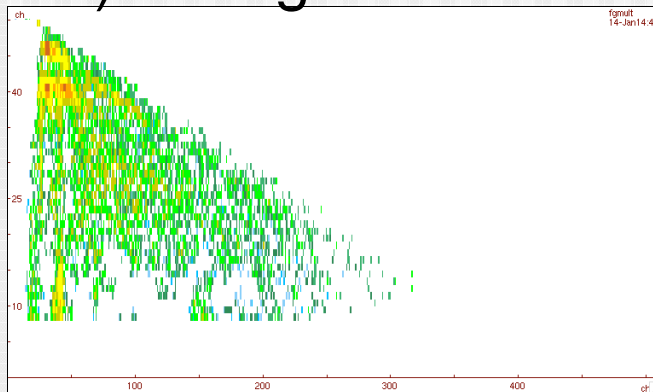
1) Raw



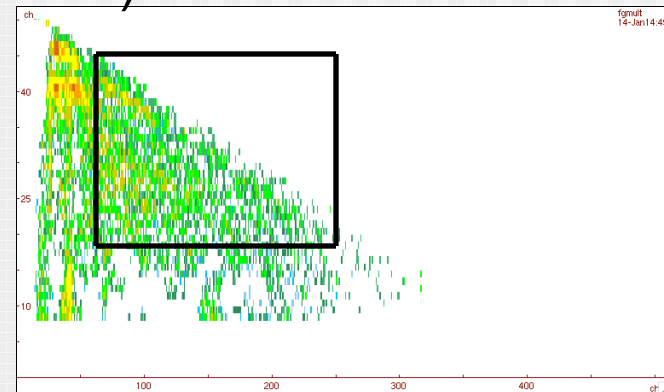
2) Unfolded



3) First generation



4) Extraction area





# Normalization

- Transformation invariance:

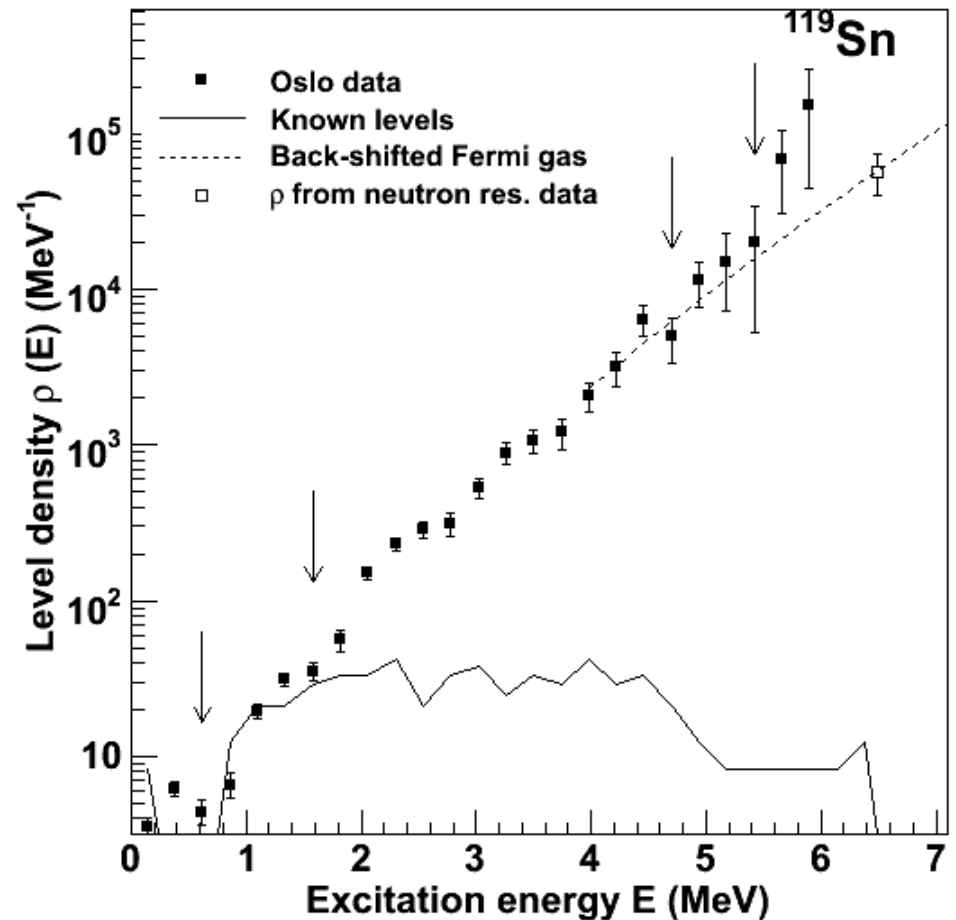
$$\tilde{\rho}(E_i - E_\gamma) = \rho(E_i - E_\gamma) \cdot A \cdot e^{\alpha(E_i - E_\gamma)}$$

$$\tilde{T}(E_\gamma) = T(E_\gamma) \cdot B \cdot e^{\alpha(E_\gamma)}$$

- Normalization of absolute value  $A$  of  $\rho$ :  
From literature value of discrete level density for low  $E_x$ .
- Normalization of common slope  $\alpha$ :  
From calculating  $\rho(B_n)$  from literature values.  
 $\rho = 1 / D$ , from  $D_0$  (s-wave resonance spacing).
- Normalization of absolute value  $B$  of  $T$ :  
From calculating  $T(B_n)$  from literature values.  
 $T = \langle \Gamma\gamma_0 \rangle / D_0$ , where  $\langle \Gamma\gamma_0 \rangle$  is average s-wave radiation width.

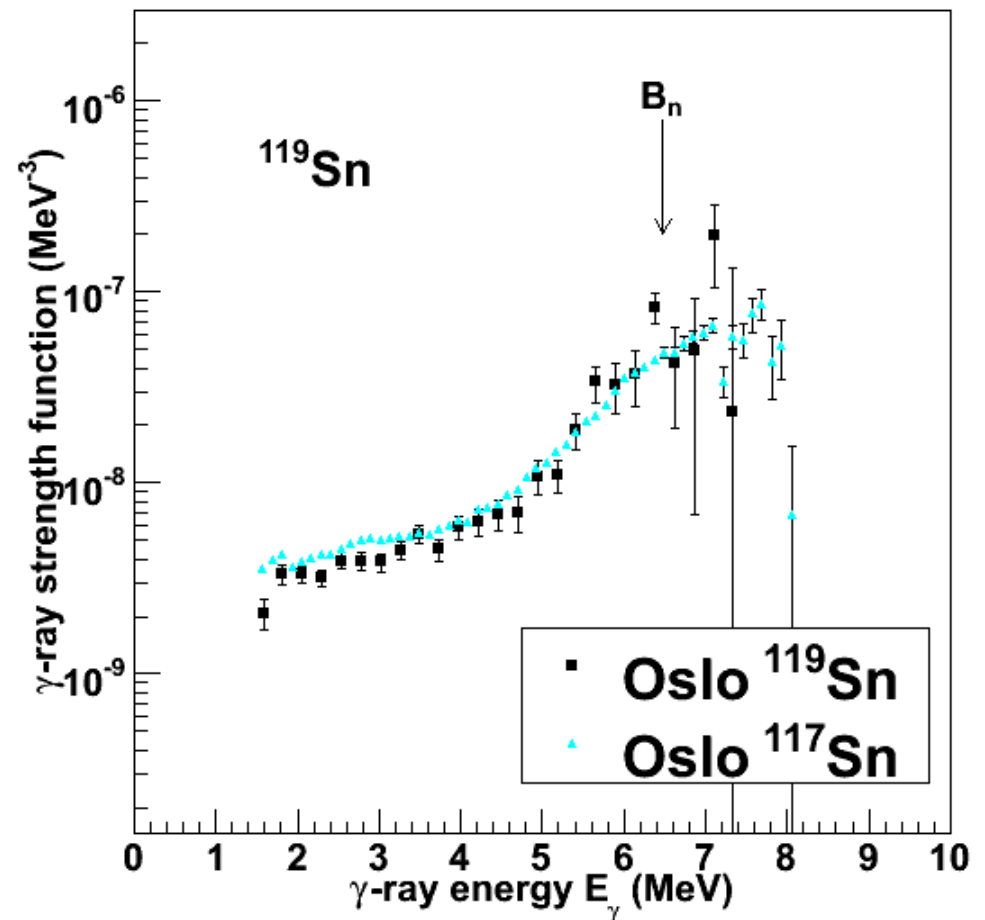
# $^{119}\text{Sn}$ level density (preliminary)

- Level density generally has a stepwise increase as a function of  $E_x$ .
- Breaking of neutrons and protons pairs.
- In tin, the steps are easily seen, since only neutron pairs are broken ( $Z = 50$ ).
- $^{119}\text{Sn}$  steps: Verification of pair-breaking published for  $^{116,117}\text{Sn}$ .  
(*PRC* **79**, 014320 (2009).)
- Future: Extraction of thermodynamic properties.



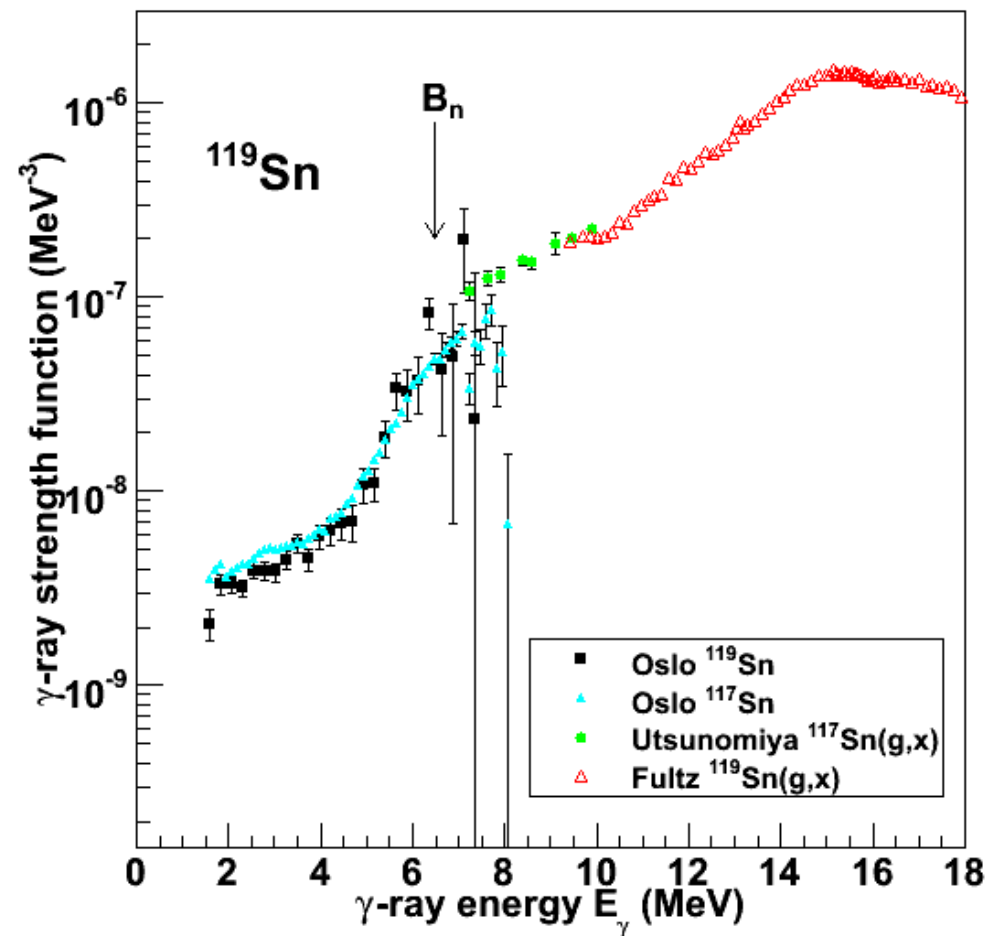
# $^{119}\text{Sn}$ $\gamma$ strength function (prelim.)

- $^{117}\text{Sn}$  strength function was recently published in PRL.
- $^{119}\text{Sn}$  is consistent with  $^{117}\text{Sn}$ .
- A sudden increase of strength at  $\approx 5$  MeV.
  - Indicates a resonance.



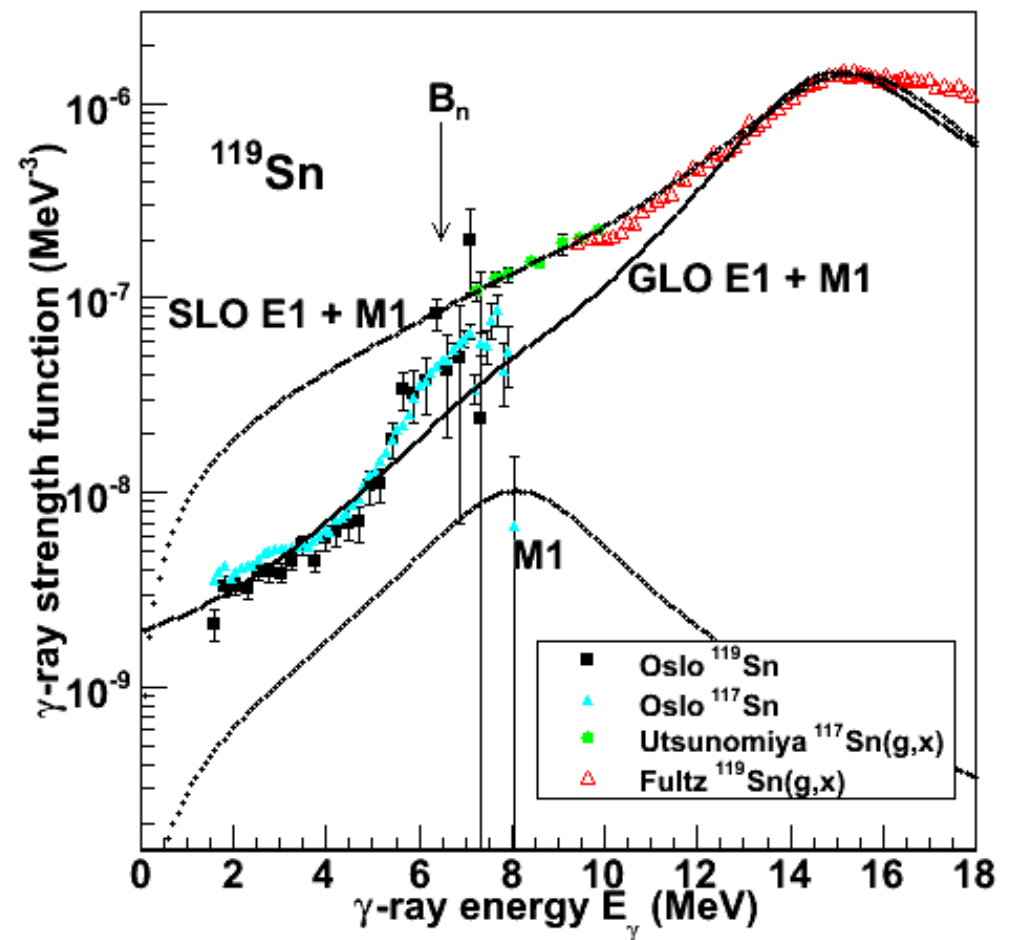
# Inclusion of $^{119}\text{Sn}$ high $E_\gamma$ data

- Add high  $E_\gamma$  data from other experiments.
- Interpretation.
- Gives the right hand side of the pygmy.



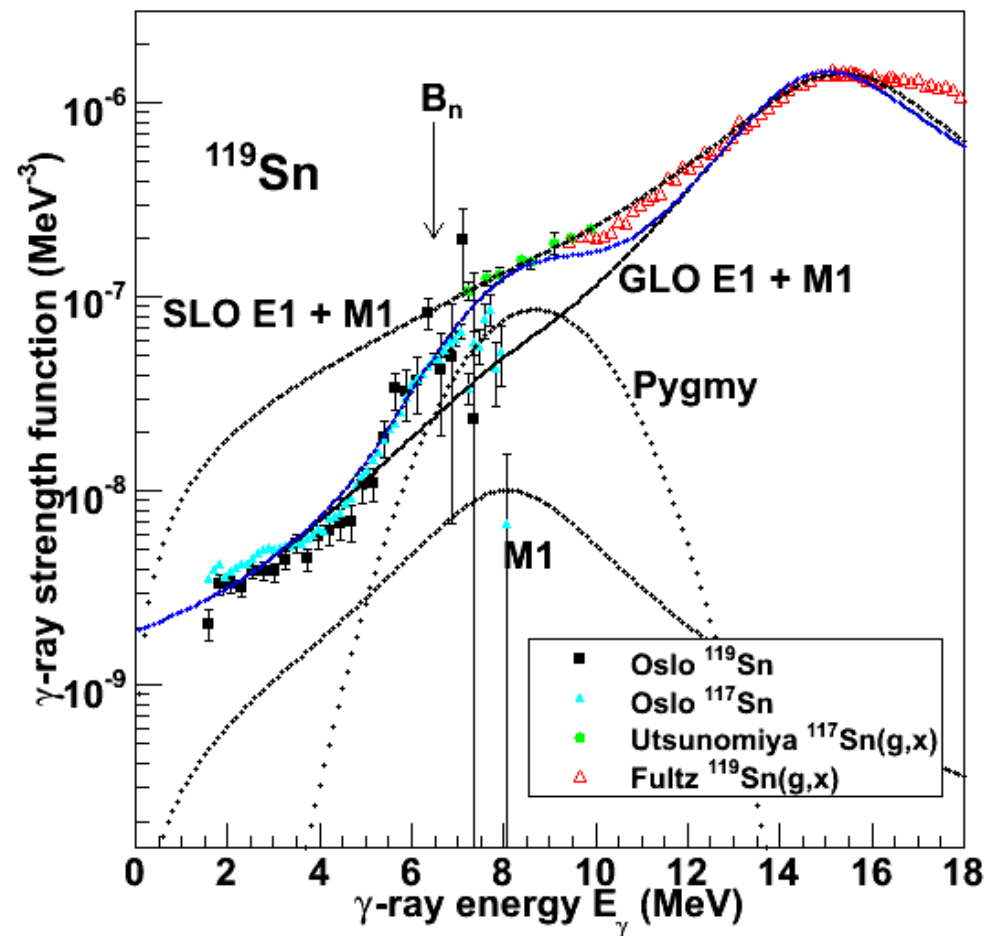
# Theoretical models for $^{119}\text{Sn}$

- SLO does not work:  
Fits to experimental values for high  $E_\gamma$ , but not low.
- GLO works:  
Fits to experimental values for both low and high  $E_\gamma$ .
- Also the case for all other investigated elements.
- Hence, a pygmy is added to a GLO baseline.



# $^{119}\text{Sn}$ pygmy fit (preliminary)

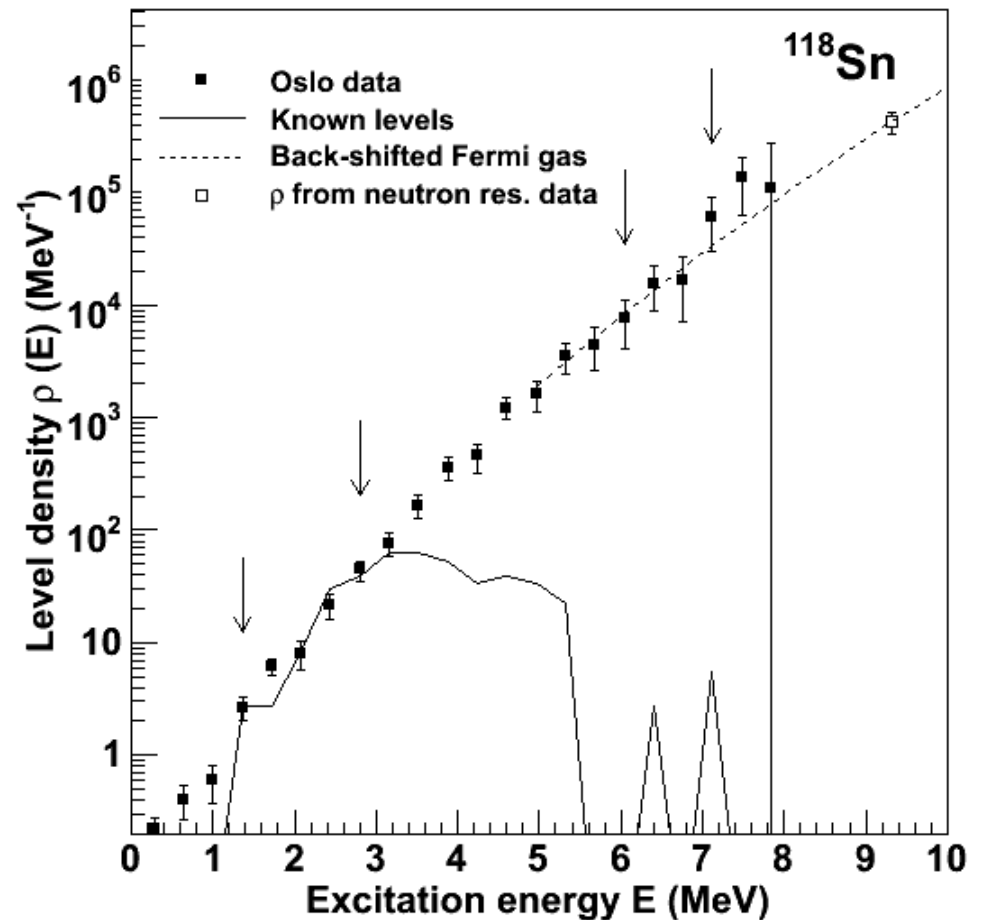
- Gaussian pygmy.
- Empirical parameters will depend on choice of baseline.
- $T = 0.35$  MeV.
- Peak  $E_{pyg} = 9.2$  MeV.
- Width  $\sigma = 1.5$  MeV.
- Normalization constant  $C_{pyg} = 4.3 \cdot 10^{-7}$  MeV<sup>-2</sup>.
- Pygmy strength: 1.8 % of Thomas-Reiche-Kuhn (TRK) sum rule.
  - Within uncertainty of  $^{117}\text{Sn}$ : 2.3(8) %.





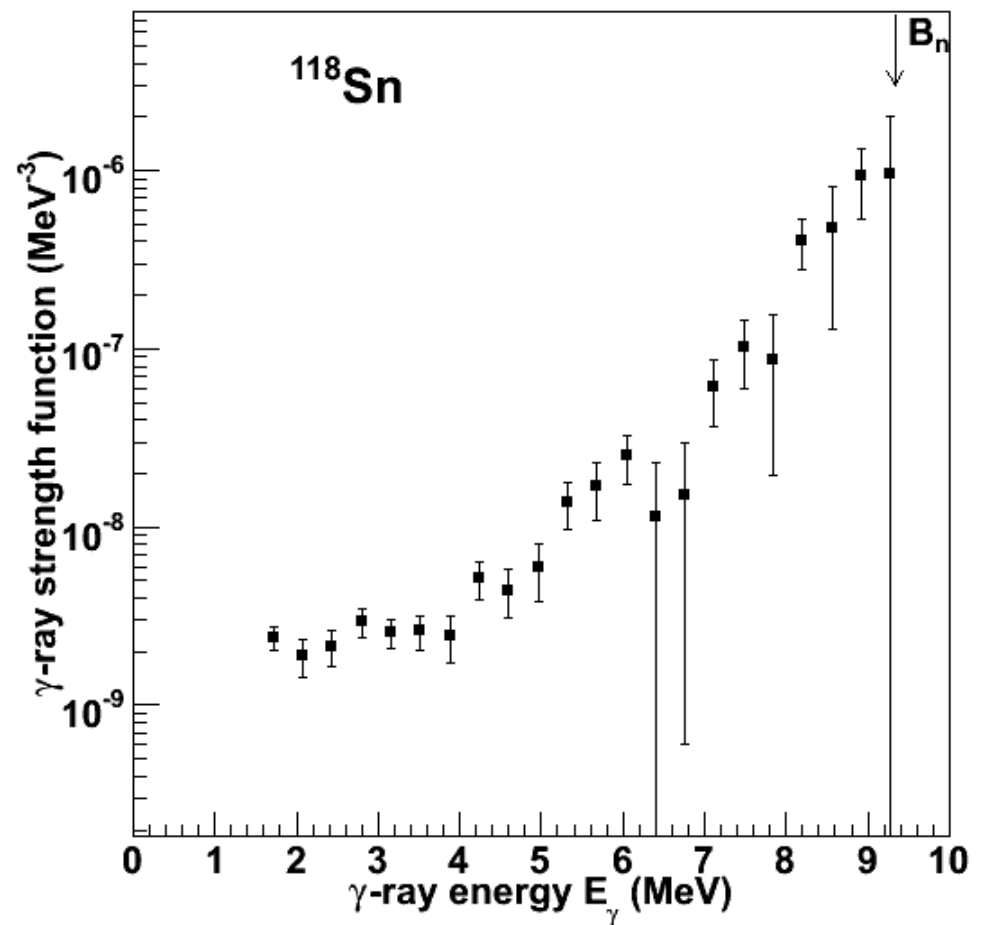
# $^{118}\text{Sn}$ level density (preliminary)

- Stepwise increase of level density was seen in  $^{116}\text{Sn}$ .
- Expected in  $^{118}\text{Sn}$ .
- Hard to verify due to less statistics.
- Small reaction cross section.



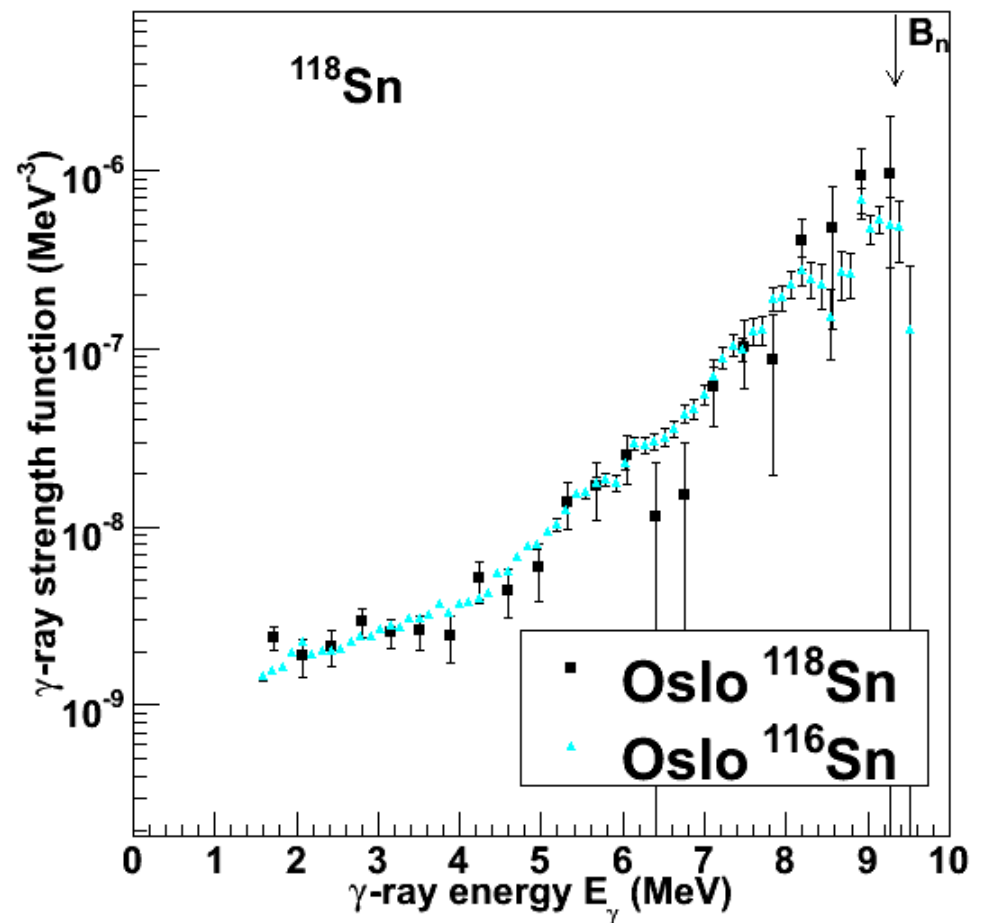
# $^{118}\text{Sn}$ strength function (preliminary)

- Same structural form as  $^{117,119}\text{Sn}$ .
- Sudden increase at  $\approx 5$  MeV.



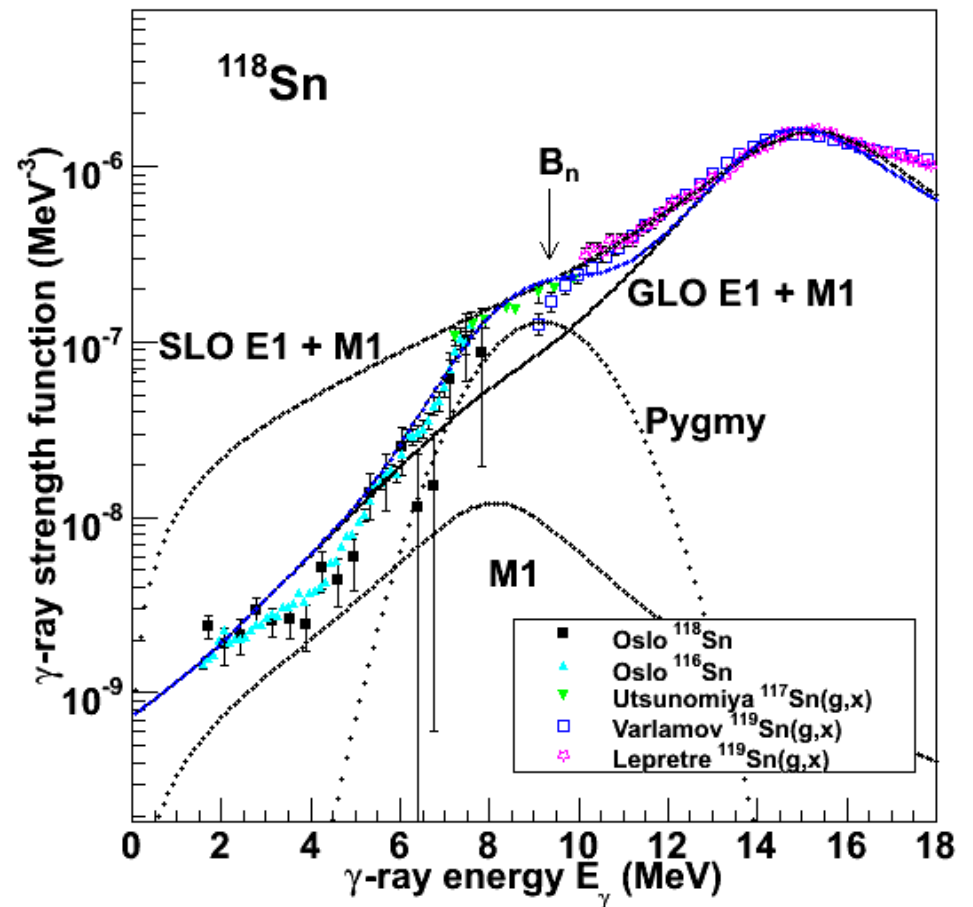
# $^{116}\text{Sn}$ strength function (preliminary)

- Lack of literature data for  $^{116}\text{Sn}$ .
- Impossible to normalize the absolute value.
- However, the  $^{118}\text{Sn}$  strength function may be used as a means to adjust  $^{116}\text{Sn}$ .
- Publishable.



# $^{116,118}\text{Sn}$ pygmy fit (preliminary)

- $T = 0.20$  MeV.
- Peak  $E_\gamma = 9.2$  MeV.
- Width  $\sigma = 1.3$  MeV.
- Normalization constant  
 $C_{pyg} = 4.0 \cdot 10^{-7}$  MeV<sup>-2</sup>.
- 2.6 % of TRK sum rule.
  - Within uncertainty of  $^{117}\text{Sn}$ : 2.3(8) %.



# Conclusions

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- As in  $^{116,117}\text{Sn}$ , a stepwise increase of level density is seen in  $^{119}\text{Sn}$ .
- As in  $^{117}\text{Sn}$ , a pygmy resonance with peak  $E_\gamma \approx 9$  MeV is seen in the strength functions of  $^{116,118,119}\text{Sn}$ .
- The position, strength and width of the pygmies indicate that they are due to the neutron-skin oscillation mode.
- TRK values for  $^{118,119}\text{Sn}$  are consistent with  $^{117}\text{Sn}$  within uncertainty.
- $^{116,117,118,119}\text{Sn}$  pygmies seem to be consistent.

